LA-UR-96-1921	CONF-96051694		
Title:	Measurement of Froton Froduction Cross Sections of ¹⁰ Be and ²⁶ Al from Elements Found in Lunar Rocks.		
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Submitted to:	Proceedings of the /th Intornational Conference on Accelerator Mass Spectrometry (held in Tucson in May 1996), to be published in Nucl. Instrum & Methode.		
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Paper 8B-1(9A-1)

Measurement of proton production cross sections of ¹⁰Be and ⁷⁶Al from elements found in lunar rocks.

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Proton production cross sections for ¹⁰Be, ²⁶Al, 7Be and 22Na produced by protons in silicon dioxide (fee oxygen) and silicon targets over a wide proton energy range, 25 - 500 MeV, have been measured using accelerator mass spectrometry or non-destructive γ -ray spectroscopy. These cross sections are important for interpreting measurements of the concentration of these nuclides made in extratomestrial materials by cosmic rays.

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1. Introduction

Cosmic rays penetrate the lunar surface and interact with the lunar rocks to produce both radionuclides and stable nuclides. Production depth profiles for long-lived radionuclides produced in lunar rocks are measured using Accelerator Mass Spectrometry (AMS). For a particular radionuclide these production depth profiles can be interpreted to give an estimate for the solar proton flux over a time period characterized by the half life of the radionuclide under study [1, 2]. This analysis is possible if and only if the all the cross sections for the interactions of all cosmic ray particles with all elements found in lunar rocks are well known. In practice, the most important cross sections needed are the proton production cross sections, because 98% of solar cosmic rays and --87% of galactic cosmic rays are protons.

The cross sections for the production of long-lived radionuclides were very difficult to measure before the development of AMS and only in recent years has significant progress been made in determining these essential cross sections [3, 4].

Oxygen and silicon are major constituents of lunar rocks. We have reported already ¹⁴C production cross sections from O and Si [4, 5] for proton energies 25 -500 MeV, and $O(p,x)^{10}$ Be from 58 - 160 MeV [6]. Here we present new measurements for the cross sections $O(p,x)^{00}$ Be, $O(p,x)^{70}$ Be, Si(p,x)¹⁰Be, Si(p,x)²⁰Al, and Si(p,x)²²Na from -30 - 500 MeV.

2. Experimental

These cross section measurements consist of four basic steps; target irradiations, short-lived radiouuclide determination using non-destructive γ -ray spectroscopy, sample preparation and AMS determination.

Target <u>irradiations</u>: Three accelerator facilities are used for the irradiations. irradiations are made for proton energies < 68 MeV at the University of California, Davis (Davis); for ~55 - 160 MeV at the Harvard Cyclotran Laboratory, Harvard University (1101.); for 200 - 500 MeV, the TriUniversities Meson Facility at the University of British Columbia (TRIUMF), At all three facilities the overall irradiation philosophy is the same, but details of the irradiation schemes vary.

At all facilities, thin target conditions are always used. This is achieved by using thin target foils and short target stacks, such that at most ~2 MeV would be lost in an individual target and < 10 MeV in the entire target stack. These conditions minimize both the loss due to protons scattering out of the stack, and the production of neutrons within the stack. SI targets are 0.0754 - 0.076 cm thick, 1.0 - 1.5 cm in diameter and weigh ~0.138 - ..0.309g, SiO₂ (for C) targets are 0.0508 - 0.056 cm thick, 1.0 - 1.5 cm in diameter and weigh ~ 0.095 - 0.214g. Catcher foils were used to minimize the effects due to receils.

At all three facilities, irradiations are designed so that the entire proton beam is intercepted by both the target stack and the monitor chamber used to get a direct measurement of the number of protons through the target. The monitor chamber is calibrated for each proton energy using a Faraday cup. Aluminum monitor folts are included at the front of the target stack and the reactions 2 Al(p,x)²⁴Na and 77 Al(p,x)⁷⁵Na used as an additional dosinetric check [5].

For most irradiations, irradiation times were designed to produce ~10⁷ ¹⁴C atoms in the control target foil and one of the eatcher foils was used for ¹⁰Be and ²⁶A1 determination by AMS. <u> γ -ray spectroscopy</u>: ⁷Be and ²¹Na are determined in all targets at HCL using non-destructive γ -ray spectroscopy.

Sample proparation: Pottowing the addition of Be and Al carriers, with the exception of the Al targets, which required no Al carrier, the targets are dissolved in addic solutions. After dissolution several steps of HCIO, furning are performed. Standard ion-chromatographic techniques are used to separate Be and Al from other species. Since the clution-of Be and Al from cation columns aliffer this technique also separates the Be from the Al. Both Al and Be are precipitated in a basic rotation as hydroxides. The Al' and Be hydroxides are then dried and converted to exides at high temperature (~ 800 C). The samples are mixed with Ag prior to AMS analysis. Successful AMS analysis of Be requires a sample preparation protocol that yields a sample relatively free of boron, an isobaric interference. With only a few exceptions this procedure produced both Al and Be samples free of interfering species. In those instances where interferences created problematic analyses repeat measurements were performed.

AMS determination: ¹⁰Be analyses for some of the silicon dioxide targets were made at the University of Arizona and the results reported [6]. All other targets were analyzed at the Lawrence Livermore National Laboratory AMS Facility (LLNL) [7] with the chemistry done at San Jose State University. Both ¹⁰Be and ²⁶Al measurements are made utilizing a dE/dX detector. For the ¹⁰Be measurements the boron background is reduced by using a thin Havar foll as the entrance window for the dE/dX detector. For most of the analyses <5% of the counts accumulated in the ¹⁰Be spectrum are attributed to boron reactions in the Havar. A correction based on a "dirty boron blank", which is propared at LLNL, is applied in all instances.

3. Results

Now cross section measurements for SI(p,x) ⁷Be, SI(p,x)¹⁰Bø, Si(p,x)²²Nn, and Si(p,x)²⁶Al, are given in Tuble 1. and now cross section measurements for $O(p,x)^7$ Be and $O(p,x)^{40}$ Be from oxygen in SiO₂ targets in Table 2; Table 2 includes corrected values for the cross sections previously reported [6].

The principal sources of error are as follows; 2% for the number of atoms/end of the target element; 5% for the number of protons through the target at all three facilities, - except for some irradiations of S1 and SiO₂ targets at FICT, where this error was 7%; and a range of values for the AMS determination depending on the number of ¹⁰Be or ²⁰Ai atoms made in the target. The irradiation conditions were optimized for ¹⁴C production, so for some target/energy combinations, too few ¹⁶Be atoms were produced for optimum AMS determinations leading to an increased error in the AMS measurement. However, even under these conditions, most cross sections were measured to better than 10%.

Results of the new cross section measurements are shown in Figures 1 - 6 for the cross sections for $Si(p,x)^{10}Be$, $Si(p,x)^{24}AI$, $Si(p,x)^{7}Be$, $Si(p,x)^{72}Na$, $O(p,x)^{10}Be$, and $O(p,x)^{7}Be$. These new measurements are compared to those in the literature and in most cases, the agreement is very good [3,8-25].

4. Discussion and Conclusions

These irrudiations are made using thin targets and small target stacks to minimize both the production of neutrons within the stack and the scattering of protons out of the stack. All irradiations are designed to include a direct measurement of the number of protons through a target stack, with monitor foils used as an additional elseck. These experimental conditions allow a simple calculation for the cross section. The majority of the new cross section values reported here, agree well with the cross sections measured over the past decade by others [3, 8, 9].

Production depth profiles of cosmogenic radionuclides measured in timer rocks can be analyzed to give estimates of the solar proton flux over a time period characterized by the half life of the radionuclide under study [1]. Accurate and precise cross section measurements of all the relevant proton production cross sections are essential for reliable estimates of the solar proton flux to be made. The progress in measuring these essential cross sections over the past decade is evident from Figures 1 - 6. Many cross sections have now been measured by more than one independent technique, with consistent results. These values can now be used with confidence to get better estimates of the solar proton flux from the ancient past to present solar cycles.

Cross sections for the production of ¹⁰Ue below ~150 MeV have only recently been measured and are similar to the estimated cross sections used by Nishilzumi (26) to interpret ¹⁰Be measured in humr rock 68815. These recent ¹⁰Be cross section measurements eliminate the possibility noted by Nishiizumi [26] that the cross sections assumed for ¹⁰He production at low energies were wrong, and support the case for a soft solar proton spectrum with relatively few high-energy protons over the last few million years.

Acknowledgments

Grant NAGW 4609 from the National Aeronautics and Space Administration provided partial support for this work at Harvard University, San Jose State University and Lawrence Livermore National Laboratory. Part of this work was performed under the auspices of the Department of Energy by Lawrence Livermore National Laboratory under contract W-7405-Eng-48. The work performed at the Los Alamos National Laboratory was supported by NASA and done under the auspices of the Department of Energy.

Proton Energy*	Si(p,x) ⁷ Be	Sl(p,x) ¹⁰ Be	Si(p,x) ²² Na	Si(p,x) ²⁶ Al
MeV	mb	mb	mb	mb
500	5.710.43	0.554±0.033	17.7±1.31	18.7:£1.09
400	4.25±0.32	0.397±0.029	17.5±1.29	20,4:1:1.18
300	3.910.3	0.262±0.0159	17.5±1.31	22.1±1.3
200	2.9+0.26	0.11450.0073	15.1±1.16	24.8 ± 1.44
157.3±1.0	1.41±0.15	0.066±0.0064	16.1±1,45	32.6±2.54
148.212.0	1.310.15	0,062:1.0.005	16.11.1.46	7.9.1±2.22
128.7±2.0				32.3±1.88
128.012.0	1.19±0.15	0.04510.0044	16.5±1.49	34.612.64
127.5±2.0			17.910.99	
127.0±2.0	1.11:0.34	0.0483.0.0038		
97 ± 3.0	0.87±0.16	0.02310.0026	16.8;11.53	36.412.76
87.413.0	0.84±0.14	0.01910.0019	18.1+1.63	38.8±3.01
77.2±:.5	0.7510.13	0.011±0.0013		43.9±3.36
67.4	0.661.0.24			
66.9	0.81:1.0.16	0.0097:t0.0012	19.911.82	47.013.57
65.6		0.005±0.0005	25.5:12.4	50.5±2.87
65.2	0.88±0.203			
61.7 <u>±</u> 4.0	0.41±0.16	0.0065;L0.0009	20.0±1.84	53,014,04
59.0			22.412.08	
58.7			22.012.09	
56.5±4.0	0.31±0.15	0.0024:10.0009	16.8:11.54	51.345.04
52.4	0.5±0.105			
50.2		0.00094.0.0002	11.6::1.09	70,244.05
44.5				67.043.95
41.0			1.1610.12	
29.8				5.994.0.38
*200-500 MeV TI	RIOMF;			

Table 1. Cross sections for the production of radionuclides from Si targets.

lower energies; with energy spread indicated, HCL; no energy spread indicated, Davis;

Proton Energy*	O(p,x) ⁷ Be	O(p,x) ¹⁰ Bc
MeV	nıb	mb
500	10.2±0.77	
400	10.3±0.77	1.18±0.068
300	11.1±0.85	0.9 52±0.05 8
200	10.3±0.77	0.712±0.046
158.9±1.0	6.82±0.63	0.56±0.048
149.9±2.0	7.11±0.66	0.51±0.057
129.9±2.0	7.3±0.69	0.·12±0.044
129.952.0	6.27::0.56	
99.9±3.0	7.57±0.74	0.44±0.047
89.9±3.0	7.25±0.68	0.33±0.038
79.9±3.5	7.66±0.73	0.27±0.032
69.9±3.5	7.69:10,75	
67.4		0.33±0.02
64.8±4	7.681.0.75	0.2110.037
59.8±4	8.44±0.83	0.18±0.022
49.9		0.038+0.0029
41.8		0.0079±0.0008
30.9		0.005240,0004

Table 2. Cross sections for the production of radionuclides from O using SiO_2 targets.

*200-500 MoV TRIUMF;

lower energies; with energy spread indicated, 11CL; no energy spread indicated, Davis;

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Figure Captions

Figure 1. $Si(p,x)^{10}Be$; errors are as shown or are smaller than the plotted symbol. Figure 2. $Sl(p,x)^{26}Al$; errors are as shown or are smaller than the plotted symbol. Figure 3. $Si(p,x)^{7}Be$; errors are as shown or are smaller than the plotted symbol. Figure 4. $Sl(p,x)^{22}Na$; errors are as shown or are smaller than the plotted symbol. Figure 5. $O(p,x)^{10}Be$; errors are as shown or are smaller than the plotted symbol. Figure 6. $O(p,x)^{7}Be$; errors are as shown or are smaller than the plotted symbol.



Piquive 1



Figure 2.









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