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TRANSPORT PHENOMENA IN A MIXTURE OF ELECTRONS AND NUCLEI

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At extremely high temperatures atoms are stripped of all or most of their electrons. The mean free path of the electrons moreover is proportional to the square of their kinetic energy. The electrons will therefore cause the ionized gas to conduct both electricity and heat quite easily. By solving the Boltzmann equation for assumed gradients of density, electric potential and temperature, we find the velocity distribution of the electrons as an expansion in Laguerre polynomials. From the first two coefficients of this expansion we find the electric and thermal conductivities. The long range of Coulomb forces leads to difficulties (divergent integrals) if one restricts the discussion to binary collisions. This is avoided by considering a shielding effect due to the rearrangement of the electrons in the neighborhood of a colliding pair.



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TRANSPORT PHENOMENA II A MIXTURE OF ELECTRONS AND NUCLEI

At extremely high temperatures atoms are stripped of all or most of their electrons. The free electrons cause the ionized gas to conduct both electricity and heat quite easily. This investigation uses the kinetic theory of gases to find expressions for the electric and thermal currents carried by the free electrons due to gradients of density, potential and temperature 1, To do this one has to determine the velocity distribution function of the electrons from the Boltzmann equation. We shall treat the problem in the approximation wherein the heavy particles shiw no deviation from the Maxwell distribution. It is not necessary to assume the electrons and the nuclei to be at the same temperature. We shall, however, not consider the resulting heat exchange and assume times independent distribution . Junctions. We confine ourselves to the linear problem with all gradients and curronts in the x direction. For the electron distribution, we try the form

$$\begin{aligned}
\phi(\mathbf{x}, \vec{\mathbf{v}}) &= \mathbf{f}(\mathbf{x}, \mathbf{v}) \begin{bmatrix} \mathbf{1} + \mathbf{v}_{\mathbf{x}} \mathbf{h}(\mathbf{v}) \end{bmatrix} \\
\mathbf{f}(\mathbf{x}, \mathbf{v}) &= \mathbf{n} \beta^3 + \frac{3/2}{2} e^{-\beta^2 \cdot \mathbf{v}^2}
\end{aligned}$$
(1)

with

$$n = n(x)$$
, $\beta = \beta(x) = \sqrt{n/2kT_{e}}$

n is the number of electrons per em².

For the nuclei of type i, we assume a distribution

$$F_{1}(xv_{i}) = N_{i}B_{i} = \pi^{-3/2} e^{-B_{i}^{2}v_{i}^{2}}$$

$$N_{i} = N_{i}(x) = \frac{B_{i}(x)}{B_{i}} = \frac{B_{i}(x)}{B_{i}} = \sqrt{\frac{M_{i}^{2}}{2kT_{i}}}$$

1) The method used here is estimily the one described by Chapman and Cowley in their book "The Mathematical Theory of Non-Uniform Gases" (Cambridge, 1939).



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(**1**a)



The Boltzmann equation can be written as

$$D(\phi) = -J_{ee}(\phi, \phi) - \sum_{i} J_{ei}(\phi, F_{i})$$
(2)

where

$$D(\phi) = v_{x} \frac{\partial \phi}{\partial x} + \frac{\Theta B}{m} \frac{\partial \phi}{\partial v_{x}}, \quad (\Theta = -l_{\mu} S \times 10^{-10} \text{ osu}) \quad (3a)$$

$$J_{eo}(\mathfrak{M}) = \iint w_{eo}(w_{e}) \left\{ \mathfrak{K}(v) \mathfrak{K}(v_{1}) - \mathfrak{K}(v_{1}) \mathfrak{K}(v_{1}) \right\} dv_{1} d\Omega \qquad (\mu)$$

 $\vec{w} = \vec{v} - \vec{v}_1$, θ is the angle between \vec{w} and \vec{w}' , and $d\Omega$ is the element of solid angle in the direction of \vec{w}_0 , $\sigma_{ee}(w\theta)$ is the cross section for electron-electron scattering. Similarly for the collisions with nuclei we have

$$J_{ei}(\mathscr{O}F_{i}) = \iint w_{i}\sigma_{ei}(w_{i}\theta) \left\{ \mathscr{O}(\vec{v})F_{i}(v_{i}) - \mathscr{O}(\vec{v})F_{i}(v_{i}') \right\} d\vec{v}_{i}d\Omega$$
(5a)

The standard procedure to obtain an approximate solution of the Boltzmann equation is to replace $D(\emptyset)$ by D(f) but leave \emptyset unchanged on the right-hand side. If we do this the left-hand side can be expressed in the following manner. We note first that:

and

$$\frac{1}{f} \frac{\delta f}{\delta x} = \frac{1}{n} \frac{dn}{dx} - \frac{1}{T_e} \frac{dT_e}{dx} \quad (3/2 - \beta^2 v^2)$$
$$\frac{1}{f} \frac{\delta f}{\delta v_x} = -2\beta^2 v_x$$

which by combination leads to

$$D(f) = \left[\frac{1}{n} \frac{dn}{dx} - \frac{eB}{kT_{e}} - \frac{1}{T_{e}} \frac{dT_{e}}{dx} (3/2 - \beta^{2}r^{2})\right] v_{x}f$$
(3b)

On the right-hand side of (La) we substitute p from (1), and obtain:

$$J_{\Theta\Theta} = \iint W_{\Theta\Theta}(W\Theta) f(v) f(v_1) \left[v_x h(v) + v_{1x} h(v_1) - v_x' h(v') - v_{1x}' h(v_1') \right] dv_1^2 d\Omega$$
(1b)

J can be greatly simplified by the observation that the heavy particles



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are much slower than the electrons so that w_i can be replaced by v_0 In addition, the only important collisions are those for which $1 - \cos \Theta \ll 1$ so that $v_i^{0} \approx v_i$ and $v^{0} \approx v_0$. We may therefore write:

 $J_{ei} = vf(v)h(v) \qquad \iint \sigma_{ei}(v\theta) F_i(v_i) (v_x - v_x^{e}) d\dot{v}_i d\Omega_{ei}(v\theta) F_i(v_i) (v_y - v_x^{e}) d\dot{v}_i d\Omega_{ei}(v\theta) (v_y - v_x^{e}) dv_i dU) (v_y - v_x^{e}) dv_i dV_{ei}(v\theta) (v_y - v_x^{e}) dv_i dV_{ei}($

The integration with respect to v_i can be carried out at once and leads to:

$$J_{\Theta i} = N_{i} \nabla f(\mathbf{v}) h(\mathbf{v}) \int \sigma_{\Theta i} (\mathbf{v} \Theta) (\mathbf{v}_{x} - \mathbf{v}_{x}) d\Omega$$

We note again that $\mathbf{v}^{9} \approx \mathbf{v}$ and express \mathbf{v}_{x}^{9} thus:
$$\mathbf{v}_{x}^{9} = \mathbf{v}(\cos \alpha \cos \Theta + \sin \alpha \sin \Theta \cos \beta)$$

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Integration with respect to Ø leads then to:

$$J_{ei} = 2\pi W_i v f(v) h(v) \int \sigma_{ei}(v\theta) (1 - \cos \theta) d(\cos \theta)$$



The cross section for collisions between electrons and nuclei with charge $Z_{\frac{1}{2}}$ (Rutherford scattering) is:

$$\sigma_{ei} = \left(\frac{Z_i e^2}{mv^2}\right)^2 (1 - \cos \theta)^{-2}$$

We can thus reduce J_{ei} further to:

$$J_{ei} = 4^{n} \lambda N_{i} \frac{v_{x}}{v^{3}} f(v) h(v)$$
(5b)

where

$$\lambda = \frac{1}{2} \int (1 - \cos \theta)^{-1} d(\cos \theta)$$
 (6)

This last integration is carried out in Appendix I. After we enter (3b), (4b), and (5b) into the Boltzmann equation (2) we are left with the problem to find h(v) from it. This can be done by expanding h(v) in terms of Laguerre polynomials²⁾ In particular we shall use the polynomials of order 3/2 and write for brevity:

$$L_{r}(\varepsilon) = L_{r}^{(3/2)}(\varepsilon)$$

2) For a discussion of the properties of these polynomials the reader is referred to Chapter 5 of SZEGO, Original Polynomials.





The L_r form a complete set and can be derived from their orthogonality relation:

$$\int_{0}^{\infty} \varepsilon^{3/2} e^{-\varepsilon} L_{r}(\varepsilon) L_{s}(\varepsilon) d\varepsilon = \frac{\Gamma(r+5/2)}{\Gamma(r+1)} \delta_{rs}$$
(7)

and $L_0(\varepsilon) = 1$. We note also that $L_1(\varepsilon) = 5/2 - \varepsilon$. We express (3b) in the form:

$$D(f) = \left[\left(\frac{1}{n} \frac{dn}{dx} - \frac{\sigma E}{kT_{\theta}} + \frac{1}{T_{\theta}} \frac{dT_{\theta}}{dx} \right) L_{0}(\beta^{2}v^{2}) - \frac{1}{T_{\theta}} \frac{dT_{\theta}}{dx} L_{1}(\beta^{2}v^{2}) \right] v_{x}f \qquad (30)$$

and substitute:

$$h(\mathbf{v}) = \sum_{s=0}^{\infty} c_s L_s(\beta^2 \mathbf{v}^2)$$
(8)

into (4) and (50).

We now multiply both sides of the Boltzmann equation by $\mathbf{v}_{\mathbf{x}} \mathbf{L}_{\mathbf{r}}(\beta^2 \mathbf{v}^2)$ and integrate over $d\mathbf{\bar{v}}_{\circ}$ On the left-hand side we obtain, using (7):

$$\Delta \delta_{\gamma r} = B \delta_{1r}$$
 (9)

where we have set:

$$A = \left(\frac{eE}{KT_{e}} - \frac{1}{n} \frac{dn}{dx} - \frac{1}{T_{e}} \frac{dT_{e}}{dx}\right) - \frac{n}{2\beta^{2}}$$
(10a)

$$B = \frac{1}{T_{\Theta}} \frac{dT_{\Theta}}{dx} - \frac{5n}{4\beta^2}$$
(10b)

On the right-hand side we obtain: $-\Sigma C_{\rm s} H_{\rm rs}$

where $H_{rs} = H_{rs}^{e} + \sum_{i} H_{rs}^{i}$

and where the H_{rs}^{e} and H_{rs}^{i} are defined as follows:

$$H_{rs}^{\circ} = \iiint w \sigma_{ee}(w \theta) f(v) f(v_1) v_x L_r(\beta^2 v^2) \Delta(v_x L_s) d \vec{v}_1 d \Omega$$
(11)

$$\Delta g = g(\vec{v}) + g(\vec{v}_1) - g(\vec{v}_1')$$

$$(12)$$

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and

$$\mathbf{H}_{rs}^{i} = \mathbf{N}_{i} \mathbf{Z}_{i}^{2} \frac{\lambda_{i}n}{3} \lambda \left(\frac{a}{m}\right)^{2} \int \frac{\mathbf{L}_{r} \mathbf{L}_{s} \mathbf{f}}{\mathbf{v}} d\mathbf{v}$$
(13)

The problem now consists in solving the set of equations

$$A\delta_{or} + B\delta_{lr} = \sum c_{s}H_{rs}$$
(14)

for the coefficients C_s. Actually we will need only the first two coefficients. We can write the current:

$$\mathbf{j} = \int \mathbf{v}_{\mathbf{x}} \boldsymbol{\theta} d\mathbf{v} = \int \mathbf{v}_{\mathbf{x}}^2 \mathbf{f} \boldsymbol{\Sigma} \mathbf{C}_{\mathbf{g}} \mathbf{L}_{\mathbf{g}} d\mathbf{v} = \frac{\mathbf{n}}{2\beta^2} \mathbf{c}_{\mathbf{g}} \mathbf{L}_{\mathbf{g}} d\mathbf{v} = \frac{\mathbf{n}}{2\beta^2} \mathbf{c}_{\mathbf{g}} \mathbf{L}_{\mathbf{g}} \mathbf{d} \mathbf{v} = \frac{\mathbf{n}}{2\beta^2} \mathbf{c}_{\mathbf{g}} \mathbf{c}_{\mathbf{g}} \mathbf{d} \mathbf{v} = \frac{\mathbf{n}}{2\beta^2} \mathbf{c}_{\mathbf{g}} \mathbf{c}_{\mathbf{g}} \mathbf{d} \mathbf{v} = \frac{\mathbf{n}}{2\beta^2} \mathbf{c}_{\mathbf{g}} \mathbf{$$

and the heat ourrent

$$q = \int \frac{mv^2}{2} v_x \phi d\vec{v} = kT_e \int (5/2 L_o - L_1) v_x^2 f 2C_g L_g d\vec{v}$$
(16)

$$q = \frac{5n}{4\beta^2} (c_0 - c_1) kT_0$$

It is convenient to take a factor:

$$\mu = \frac{8\sqrt{\pi}}{3} \left(\frac{ne^2}{m}\right)^2 \lambda \beta \qquad (17)$$

out of the matrix elements and to write:

$$H_{rs} = \mu h_{rs} \tag{18}$$

 C_0 and C_1 can be written formally in terms of the dimensionless matrix elements h_{rs} as follows:



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However, these determinants are infinite and the following limiting process is used to get convergent results. We cut both numerator and denominator determinants off beyond the row and column carrying the index n, take the ratio, and repeat with larger n. To carry this through we write:

$$D^{(n)} = \begin{pmatrix} h_{00}h_{01} & \cdots & h_{0n} \\ h_{n0}h_{n1} & \cdots & h_{nn} \end{pmatrix}$$
(20)

we also use the minors D_{ik}⁽ⁿ⁾ which are obtained by deleting the ith row and the kth column.

Then we form

$$R_{ik} = \frac{D_{ik}}{D} = \lim_{n \to \infty} \frac{D_{ik}}{D^{(n)}}$$
(21)





By substitution into (19) we obtain:

$$c_{0} = (AR_{00} - BR_{10})\mu^{-1}, \qquad c_{1} = (-AR_{01} + BR_{11})\mu^{-1}$$
 (22)

The limiting process (21) can be carried through by means of a theorem on determinants by Sylvester³. From this theorem we obtain the relation:

$$\begin{vmatrix} D^{(n)} & D^{(n+1)} \\ n+1,k \\ D^{(n+1)}_{i,n+1} & D^{(n+1)}_{ik} \end{vmatrix} = D^{(n+1)} D^{(n)}_{ik}$$

so that

$$\frac{D_{ik}^{(n+1)}}{D^{(n+1)}} = \frac{D_{ik}^{(n)}}{D^{(n)}} + \frac{D_{i,n+1}^{(n+1)} D_{n+1,k}^{(n+1)}}{D^{(n)} D^{(n+1)}}$$

and further:

$$R_{ik} = \frac{D_{ik}^{(1)}}{D^{(1)}} + \frac{D_{i2}^{(2)}D_{2k}^{(2)}}{D^{(1)}D^{(2)}} + \frac{D_{13}^{(3)}D_{3k}^{(3)}}{D^{(2)}D^{(3)}} + \cdots$$
(23)

It is interesting to note a simplification which is introduced if one imposes the condition j = 0 or its equivalent $C_0 = 0_0$. In this case we can eliminate A from (22) and obtain:

$$C_1 = \mu^{-1} \frac{D_{11}D_{00} - D_{10}D_{01}}{D D_{00}} B = \mu^{-1} \frac{D_{11}O1}{D_{00}} B$$
 (24)

by another application of Sylvester's theorem. We therefore do not need the zero row and column of our matrix if we only want the heat conductivity. If, however, we also want to know the value of A, that is the electric field in the case j = 0 or if we want the coefficients in the more general case when $j \neq 0$ we have to use the complete matrix.

3) See e.g. Kowalewski, THEASTEDER DETRAINANTEN, theorem 30.

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By combining (10a), (10b), (17) and (22) we are led to the equations:

$$J = \frac{3\lambda^{-1}}{l_{L}\sqrt{2\pi}} \left(\frac{m}{e^{2}}\right)^{2} \left(\frac{kT_{e}}{m}\right)^{5/2} \left[\left(\frac{eE}{kT_{e}} - \frac{1}{n} \frac{dn}{dx} - \frac{1}{T_{e}} \frac{dT_{e}}{dx}\right) R_{00} - \frac{5}{2} \frac{1}{T_{e}} \frac{dT_{e}}{dx} R_{01}\right]$$
(25)

$$q = \frac{75\lambda^{-1}}{16\sqrt{2\pi}} \left(\frac{m}{e^{2}}\right)^{2} \left(\frac{kT_{e}}{m}\right)^{5/2} kT_{e} \left[-\frac{1}{T_{e}} \frac{dT_{e}}{dx} \left(R_{01} + R_{11}\right) + \frac{2}{5} \left(\frac{eE}{kT_{e}} - \frac{1}{n} \frac{dn}{dx} - \frac{1}{T_{e}} \frac{dT_{e}}{dx}\right)$$
(26)

$$\left(\frac{R_{00} + R_{01}}{2}\right)$$
(26)

we are particularly interested in the case j = 0 where we find:

$$\frac{eE}{kT_e} = \frac{1}{n} \frac{dn}{dx} + \frac{1}{T_e} \frac{dT_e}{dx} \left(1 + \frac{5}{2} \frac{R_{01}}{R_{00}}\right)$$
(27)

$$q = -\frac{75\lambda^{-1}}{16\sqrt{2\pi}} \left(\frac{m}{\sigma^2}\right)^2 \left(\frac{kT_0}{m}\right)^{5/2} k \left(\frac{R_{00}R_{11} - R_{01}^2}{R_{00}}\right) \frac{dT_0}{dx}$$
(28)

That is we get for the heat conductivity:

$$k = \frac{75\lambda^{-1}}{16\sqrt{2\pi}} \left(\frac{R_{00}R_{11} - R_{01}^2}{R_{00}} \right) k_0 \left(\frac{e^2}{m_0^2} \right)^{-2} \left(\frac{kT_e}{m_0^2} \right)^{5/2}$$
(29)

In this formula c was introduced to put the dimensions of k into evidence. For the value of λ see equation (46).

The integrations (11) and (13) are carried out in Appendix II. The matrix h_{rs} is seen to depend on an effective nuclear charge:

$$z = \frac{\sum N_i z_i^2}{n}$$
(30)

The computations were carried through for Z value of 1,2, 2.5, 3 and ∞ . The case Z = ∞ means that the term h_{rs}^{c} resulting from electron-electron scattering was neglected. This case is very important because the Boltzmann equation can





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also be solved in a closed form, so that we are able to check our theory. The table below shows the first 4 terms and their totals of the series for R_{00} , R_{01} and R_{11} . Obviously we have $R_{10} = R_{01}$ on account of the symmetry. It also shows the important combinations, $1 + \frac{5}{2} \frac{R_{01}}{R_{00}}$ and $\frac{75}{16\sqrt{2\pi}} \frac{R_{00}R_{11} - R_{01}}{R_{00}}$ which

2 ->	1	2	2.5	3	ω
Terms of R ₀₀ 1st 2nd 3rd 4th total	1.9320 .0179 .0117 .0041 1.9657	1.1590 .0015 .0023 .0006 1.1634	.9748 .0002 .0011 .0002 .9763	. 8430 .0000 .0005 .0001 .8436	$3.2500 \ z^{-1}$.1406 z^{-1} .0039 z^{-1} .0005 z^{-1} $3.3950 \ z^{-1}$
Terms of R _{O1} 1st 2nd 3rd 4th total	.6213 0668 .0053 0015 .5583	•4393 -•0192 •0011 ••0005 •4207	∘3832 -∘0063 ∘0006 -∘0003 ∘377 ²	•3398 •0027 •0004 ••0001 •3428	1.5000 2-1 .5625 2-1 .0234 2-1 .0014 2-1 2.0377 2-1
Terms of R ₁₁ lst 2nd 3rd 4th total	.4142 .2194 .0024 .0005 .6665	•2929 •2504 •0006 •0004 •5443	。2555 。2440 。0004 。0003 。5002	.2265 .2360 .0003 .0002 .4630	$\begin{array}{r} 1.0000 \ z^{-1} \\ 2.2500 \ z^{-1} \\ .1406 \ z^{-1} \\ .0039 \ z^{-1} \\ 3.3945 \ z^{-1} \end{array}$
$1 + \frac{5}{2} \frac{R_{01}}{R_{00}}$	1.710	1.904	1.966	2.016	2.5005
$\frac{75}{16\sqrt{2\pi}} \left(\frac{R_{00}R_{11} - R_{01}^2}{R_{00}} \right)$	1-100 19498	-10192 17334	- 1.041 . 6629.	- - 952 , 6053	-6+3786 z ⁻¹ 4.0607

occur in (27) and (29).

For large Z we obtain by combining (35) and (56):

 $\frac{1}{n} \frac{dn}{dx} - \frac{eE}{kT_e} + (\beta^2 v^2 - \frac{3}{2}) \frac{1}{T_e} \frac{dT_e}{dx} = -4\pi\lambda nZ \left(\frac{e^2}{m}\right)^2 \frac{h(v)}{v^2}$ (31)



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For the electron and heat currents we have:

$$\mathbf{j} = \int \mathbf{v}_{\mathbf{x}} \mathbf{d} \mathbf{d} = \int \mathbf{v}_{\mathbf{x}}^{2} \mathbf{h} \mathbf{f} \mathbf{d} \mathbf{v} = \frac{1}{3} \int \mathbf{v}^{2} \mathbf{h} \mathbf{f} \mathbf{d} \mathbf{v}$$
(32)

$$q = \int \frac{m}{2} v^2 v_x \rho d\vec{v} = \frac{1}{3} \frac{m}{2} \int v^{\mu} h f d\vec{v}$$
(33)

To carry these integrations through we need: .

$$\int v^{5} f d\vec{v} = 2n\pi^{-1/2} \beta^{-5} \int e^{3} e^{-\varepsilon} d\varepsilon = 2n\pi^{-1/2} \beta^{-5} \cdot 31$$
(34)

and similarly

$$\int v^7 f d\vec{v} = 2n \pi^{-1/2} \beta^{-7} \mu i$$
(35)

$$\int v^{9} f dv = 2n \pi^{-1/2} \beta^{-9} 5$$
 (36)

We substitute h(v) from (31) into (32) and (33) and use the integrals (34), (35), and (36) to obtain:

$$\mathbf{j} = \frac{8}{\pi\sqrt{2\pi}} \lambda^{-1} z^{-1} \left(\frac{e^2}{m}\right)^{-2} \left(\frac{kT_e}{m}\right)^{5/2} \left(\frac{eE}{kT_e} - \frac{1}{n} \frac{dn}{dx} - \frac{5}{2} \frac{1}{T_e} \frac{dT_e}{dx}\right)$$
(37)

$$q = \frac{32}{\pi\sqrt{2\pi}} \lambda^{-1} z^{-1} \left(\frac{e^2}{m}\right)^{-2} \left(\frac{kT_e}{m}\right)^{5/2} kT_e \left(\frac{eE}{kT_e} - \frac{1}{n} \frac{dn}{dx} - \frac{7}{2} \frac{1}{T_e} \frac{dT_e}{dx}\right)$$
(38)

Comparing these equations with (25) and (26) we see that the following equalities should exist:

$$\frac{3}{4}R_{00} = \frac{8}{\pi} z^{-1}; \qquad \frac{3}{4}\left(R_{00} + \frac{5}{2}R_{01}\right) = \frac{20}{\pi} z^{-1}$$

$$\frac{15}{8}\left(R_{00} + R_{01}\right) = \frac{32}{\pi} z^{-1}; \qquad \frac{75}{16}\left(R_{01} + R_{11} + \frac{2}{5}R_{00} + \frac{2}{5}R_{01}\right) = \frac{112}{\pi} z^{-1}$$

That is we should have:

$$R_{00} = R_{11} = \frac{32}{3\pi} z^{-1} = 3.395,305 z^{-1}, R_{01} = \frac{32}{5\pi} z^{-1} = 2.03718 z^{-1}$$

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and for the derived quantities:

$$1 + \frac{5}{2} \frac{R_{01}}{R_{00}} = 2.5; \quad \frac{75}{16\sqrt{2\pi}}, \quad \frac{R_{00}R_{11} - R_{01}^2}{R_{00}} = \frac{32}{\pi\sqrt{2\pi}} z^{-1} = \frac{4.0635}{6.3830} z^{-1}$$

These values agree very well with those listed in the proceeding table, so that we can be quite confident of the validity of our general methods

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APPENDIX I. CALCULATION OF λ

The integral λ given by (6)

$$\lambda = \frac{1}{2} \int_{\theta_1}^{\theta_2} (1 - \cos \theta)^{-1} d \cos \theta = \frac{1}{2} \ln(1 - \cos \theta) \Big|_{\theta_1}^{\theta_2}$$

diverges if one uses $\Theta_1 = 0$ for the lower limit. The reason is that the kinetic theory, as it is used, restricts itself to the consideration of encounters between only two particles at a time. The Coulomb force law is, however, of such a nature that the possibility of interactions between more than two particles must not be excluded. Another, less catastrophic, difficulty arises out of the uncertainty principle which excludes the possibility of head-on collisions, because one has te consider an electron as being spread out over a region of the order of magnitude of its de Broglie wave Tength. To remedy the situation we, first of all, express λ in terms of the collision parameter p.

$$\lambda = \frac{1}{2} \ell_{m} \left[1 + \left(\frac{mv^{2}}{2e^{2}} \right)^{2} p^{2} \right] \Big|_{p_{1}}^{p_{2}}$$
(39)

The lower limit is, according to the uncertainty principle, the de Broglie wave length. That is we have:

$$P_1 \approx \frac{t_1}{m_V}$$
 (40)

The expression:

$$\left(\frac{mv^2}{2\sigma^2}\right) \mathbf{p}_1 = \frac{137}{2} \cdot \frac{\mathbf{v}}{\sigma}$$

is in our case considerably larger than one so that we can leave the one in front out. The same is obviously true at the upper limit so that we can write:

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(44)



We chose the upper limit by excluding collisions which last longer than the time during which the electron gas would be able to rearrange its density distribution so as to give a shielding effect. The rate at which this will take place is determined by the frequency of the plasma vibrations⁽⁴⁾.

$$\omega = \sqrt{\frac{\mu \pi_{\rm mo}^2}{m}} \tag{42}$$

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The collision parameter will thus be given to the right order magnitude by the relation

$$P_2 \approx \frac{v}{\omega}$$
 (43)

Thus we obtain:

$$\lambda = l_{m} \frac{mv^{2}}{\hbar \omega} = l_{m} \frac{3kT_{e}}{\hbar \omega}$$
(44)

or, after introducing (42) and rearranging to put dimensions into evidence:

$$\lambda = \frac{1}{2} \left[\frac{9}{4\pi} \left(\frac{\hbar c}{e^2} \right) \left(\frac{kT_0}{mc^2} \right)^2 n^{-1} \left(\frac{\hbar}{mc} \right)^{-3} \right]$$
(45)

For convenience we introduce Avogadro's number N and obtain:

$$\lambda = 10.881 + \ln\left(\frac{kT_0}{mc^2}\right) - \ln\left(\frac{n}{N}\right)$$
(46)

Because of the slow variation with Te it will usually be sufficient to use an average Te in this expression.

4) See Cobine, Gaseous Conductors (1941) p. 132.



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APPENDIX II. CALCULATION OF THE MATRIX ELEMENTS

In carrying through the integration (11) one can replace $v_x \Delta(v_x L_s)$ by $\frac{1}{3} \stackrel{?}{\neq} \Delta(\stackrel{?}{\neq} L_s)$ because the rest of the integrand is spherically symmetrical in the velocities. Let us furthermore express the velocities in terms of the velocity $\stackrel{?}{u}$ of the center of mass and the relative velocities $\stackrel{?}{w}$ and $\stackrel{?}{w}^{e}$ of the two particles before and after collision:

$$\vec{v} = \vec{u} + \frac{1}{2}\vec{w}, \vec{v}_1 = \vec{u} - \frac{1}{2}\vec{w}, \vec{v}_2 = \vec{u} + \frac{1}{2}\vec{v}' \quad \vec{v}_1' = \vec{u} - \frac{1}{2}\vec{w}'$$
 (47)

We shall collect these four equations symbolically in one and write:

$$\vec{v}_{i} = \vec{u} + \frac{1}{2} \vec{w}_{i} \quad (i = 1...4)$$
(48)

We further introduce the angle G; by

$$\vec{w} \cdot \vec{w}_{i} = w^{2} \cos \Theta_{i}$$
 (49)

That means that Θ_{i} assumes the values $0, \pi, \Theta, \pi - \Theta_{o}$ The Jacobian of the transformation (47) has the absolute value one so that $d\vec{v}d\vec{v}_{1} = d\vec{u}d\vec{v}$. Now let: $W_{i} = (f(x)f(x_{i})(\vec{x}_{i},\vec{x}_{i}))^{-\beta} (xv^{2}+yv^{2}_{i})$

$$\mathbf{u}_{\mathbf{i}} = \int \mathbf{f}(\mathbf{v}) \mathbf{f}(\mathbf{v}_{\mathbf{i}}) \left(\mathbf{v} \cdot \mathbf{v}_{\mathbf{i}} \right) e^{-\beta^2 (\mathbf{x} \mathbf{v}^2 + \mathbf{y} \mathbf{v}_{\mathbf{i}}^2)} d\mathbf{u}^{+}$$
(50)

where

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$$x = \frac{\xi}{1 - \xi}, \quad y = \frac{\gamma}{1 - \eta}$$
 (51)

Then, considering the generating function of the Laguerre polynomials:

$$(1 - \xi)^{-5/2} \circ^{-xc} = \sum_{r} \xi^{r} L_{r}(s)$$

$$(52)$$

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we can write:

$$\sum_{\mathbf{r}} \sum_{\mathbf{s}} \xi^{\mathbf{r}} \eta^{\mathbf{s}} H_{\mathbf{r}s}^{\mathbf{e}} = \frac{1}{3} (1 - \xi)^{-5/2} (1 - \eta)^{-5/2} \iint w\sigma_{\mathbf{e}\mathbf{e}}(w\theta) (M_{1} + M_{2} - M_{3} - M_{4}) dwdn.$$
(53)

 H_{rs}^{6} appears thus as a coefficient in expanding the expression (53) in powers of \S and η . Our next step is therefore to determine the integrals $M_{i^{0}}$ Introducing (1a) we obtain:

$$\mathbf{M}_{i} = n^{2} \left(\frac{\beta^{2}}{n}\right)^{3} \int \mathbf{e}^{-\beta^{2}} \left[(1+x) \mathbf{v}^{2} + \mathbf{v}_{1}^{2} + \mathbf{y} \mathbf{v}_{i}^{2} \right] \left(\mathbf{v} \cdot \mathbf{v}_{i} \right) d\mathbf{u}$$
(54)

Introducing the substitution (47) we rewrite:

$$(1 + x)v^{2} + v_{1}^{2} + yv_{1}^{2} = (u^{2} + \frac{1}{4}w^{2})(2+x+y) + \vec{u} \cdot (x\vec{w} + y\vec{w}_{1})$$
$$= (2 + x + y)g^{2} + jw^{2}$$
(55)

where:

$$\vec{g} = u + \frac{x\vec{w} + y\vec{w}_1}{2(2+x+y)}$$
 (56)

$$j = \frac{2(1 + x + y) + xy(1 - \cos \Theta_{j})}{2(2 + x + y)}$$
(57)

(56) is simply a change of origin so that we have $d\vec{u} = d\vec{g}$. We express \vec{v}_i in terms of \vec{g} as

$$\vec{v}_1 = \vec{g} + \vec{g}_1$$
(58)

with:

$$\vec{b}_1 = \frac{(2+x)\vec{w}_1 - x\vec{v}}{2(2+x+y)}$$
 (59)

Similarly we have:

 $\mathbf{v} = \mathbf{\dot{g}} + \mathbf{\dot{g}}_0 \tag{60}$

with:

$$\frac{(2 + y)\vec{v} - y\vec{v}_1}{a(2 + x + y)}$$
(61)

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Entering (55), (58) and (60) into (54), we can carry out the integration and obtain:

$$M_{1} = \pi^{-3/2} n^{2} \beta (2 + x + y)^{-5/2} e^{-j \beta^{2} w^{2}} \left(\frac{3}{2} + \beta^{2} (2 + x + y) \vec{g}_{0} \cdot \vec{g}_{1} \right)$$
(62)

where $\vec{g}_0 \cdot \vec{g}_1$ can be obtained from (59) and (61)

$$\vec{B}_0 \circ \vec{B}_1 = \frac{-(x + y + xy) + (2 + x + y + xy) \cos \Theta_1}{2(2 + x + y)^2}$$
 (63)

In order to carry through the integration (53) we set:

$$A = \frac{2}{2} n^{-3/2} \frac{2}{n} \beta (2 + x + y)^{-5/2}$$

$$B = \frac{x + y + xy}{3(2 + x + y)} \beta^{-2}$$

$$G = \frac{2 + x + y + xy}{3(2 + x + y)} \beta^{-2}$$

$$D = \frac{2 + 2x + 2y + xy}{2(2 + x + y)} \beta^{-2}$$

$$B = \frac{xy}{2(2 + x + y)} \beta^{-2}$$
(64)

Thus we can write:

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$$\mathbf{M}_{\mathbf{i}} = \mathbb{A}(\mathbf{1} - \mathbf{B}\mathbf{w}^{2} + \mathbf{C}\cos\boldsymbol{\Theta}_{\mathbf{i}}\mathbf{w}^{2}) \mathbf{e}^{-(\mathbf{D} - \mathbf{E}\cos\boldsymbol{\Theta}_{\mathbf{i}})\mathbf{w}^{2}}$$
(65)

Now we form $\Delta M_1 = M_1 + M_2 - M_3 - M_4$ and expand in powers of $v = \cos \theta - 1$. Actually, we will need only the linear term of the expansion because the scattering pross-section is proportional to v^{-2} so that the quadratic and higher terms give small contribution to the integral as compared with the linear terms

Entering the proper
$$\mathfrak{S}_{1}$$
 values we obtain:

$$M_{1} = \Lambda(1 - Bw^{2} + Cw^{2}) e^{-(D-E)w^{2}}$$

$$M_{3} = \Lambda(1 - Bw^{2} + Cwcor \theta e^{-(D-E)cos \theta})w^{2}$$
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and by subtracting

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$$\mathbf{M}_{1} - \mathbf{M}_{3} = A e^{-(D-E)w^{2}} \left[(1 - Bw^{2}) (1 - e^{Evw^{2}}) + Cw^{2} (1 - (v + 1)e^{Evw^{2}}) \right]$$

= $A e^{-(D-E)w^{2}} \left[(1 - Bw^{2})Ev^{2} + Cw^{2} (1 + Ew^{2}) + O(v^{2}) \right]$ (67)

 $M_2 - M_1$ is obtained by simply changing the signs of C and E_o The crosssection of e - e scattering is:

$$\sigma_{ee} = \left(\frac{2e^2}{m_e^2}\right)^2 v^{-2}$$
(68)

We now determine the integral:

$$\int w\sigma_{ee}(w\theta) (M_1 - M_3) d\vec{w} d\Omega$$

$$= (4\pi)^2 \lambda \left(\frac{2e^2}{m}\right)^2 \Lambda \int e^{-(D-E)w^2} \left[(E+C) - E(B-C)w^2 \right] wdw \qquad (69)$$

$$= 32\pi^2 \lambda \left(\frac{e^2}{m}\right)^2 \left[\frac{E+C}{D-E} - \frac{E(B-C)}{(D-E)^2} \right]$$

and

$$\int w\sigma_{ee}(\mathbf{M}_{2} - \mathbf{M}_{1}) d\vec{\mathbf{w}} \mathbf{i} \Omega = 32\pi^{2} \lambda \left(\frac{\mathbf{e}^{2}}{m} \mathbf{A} \left[-\frac{\mathbf{E} + \mathbf{C}}{\mathbf{D} + \mathbf{E}} + \frac{\mathbf{E}(\mathbf{B} + \mathbf{C})}{(\mathbf{D} + \mathbf{E})^{2}} \right]$$
(70)

so that;

$$\int w\sigma_{ee} \Delta M_{i} d\vec{w} d\Omega = 64\pi^{2} \lambda \left(\frac{e^{2}}{m}\right)^{2} \Delta B \frac{D^{2}E + 2D^{2}C - E^{3} - 2BDE}{(D^{2} - E^{2})^{2}}$$
(71)

Now we have to express (71) as a function of
$$\xi$$
 and γ . If we set
 $a = (1 - \xi)^{-1}(1 - \eta)^{-1}$ we find:
 $2 + x + y = (2 - \xi - \eta)a$
 $x + y + xy = (\xi + \eta - \xi \eta)a$
 $2 + x + y + xy = (2 - \xi - \eta + \xi \eta)a$
 $\ell + 2x + 2y + xy = (2 - \xi \eta)a$
 \vdots
 \vdots
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and entering this into (64):

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$$A = \frac{3}{2} \pi^{-3/2} n^2 \beta (2 - \xi - \eta)^{-5/2} a^{-5/2}$$

$$B = \frac{1}{3} \frac{\xi + \eta - \xi \eta}{2 - \xi - \eta} \beta^2$$

$$C = \frac{1}{3} \frac{2 - \xi - \eta + \xi \eta}{2 - \xi - \eta} \beta^2$$

$$D = \frac{1}{2} \frac{2 - \xi - \eta}{2 - \xi - \eta} \beta^2$$

$$E = \frac{1}{2} \frac{\xi \eta}{2 - \xi - \eta} \beta^2$$

We enter these expressions into (71) and multiply according to (53) by

$$\frac{1}{3}(1-\xi)^{-5/2}(1-\eta)^{-5/2} = \frac{1}{3} a^{5/2} \text{ and get:}$$

$$\sum_{r} \sum_{s} \xi^{r} \eta^{s} H_{rs}^{\circ} = \mu \sqrt{2} \frac{\xi \eta \left(1-\frac{1}{2}\left(\xi+\eta\right)-\frac{1}{8}\left(\xi\eta\right)+\frac{1}{12}\left(\xi\eta\right)\left(\xi+\eta\right)-\frac{3}{8}\left(\xi\eta\right)^{2}\right)}{\left(1-\frac{1}{2}\left(\xi+\eta\right)\right)^{5/2}\left(1-\xi\eta\right)^{2}} (\eta^{1})$$

where μ is defined by (17). We can see immediately that all elements in the zero rew and zero column are zero. By expanding the expression (74) in powers of ξ and γ , we obtain the symmetrical matrix:

$$h_{rs}^{\circ} = \sqrt{2} \qquad \begin{cases} 0 & 0 & 0 & 0 & 0 \\ 1 & 3/2^2 & 15/2^5 & 35/2^9 & \dots \\ & 45/2^4 & 309/2^7 & 885/2^9 & \dots \\ & 5657/2^{10} & 20349/2^{12} & \dots \\ & 149749/2^{14} & \dots \\ & & 149749/2^{14} & \dots \\ \end{cases}$$
(75)

The integration (13) requires considerably less labor. The angular integration gives a factor 4^{π} and if we further set $\beta^2 v^2 = \epsilon$ and use (1a) we have:

$$H_{rs}^{1} = \mu \frac{N_{1}Z_{r}^{2}}{L_{L}} \int_{0}^{\infty} L_{r}(e L_{s}) e e^{-e} d e$$
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(76)

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as before, we make use of the generating function, and write:

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$$\sum \sum \xi^{r} \eta^{s} h_{rs}^{i} = \frac{N_{i} Z_{i}^{2}}{n} (1 - \xi)^{-5/2} (1 - \eta)^{-5/2} \int_{0}^{\infty} e^{-(x+y+1)\varepsilon} d\varepsilon$$

$$= \frac{N_{i} Z_{i}^{2}}{n} (1 - \xi)^{-3/2} (1 - \eta)^{-3/2} (1 - \xi\eta)^{-1}$$
(77)

By expanding this in powers of f and η , we obtain the symmetrical matrix :

$$h_{rs}^{1} = \frac{N_{1}Z_{1}^{2}}{n} \begin{cases} 1 & \frac{3}{2} & \frac{15}{2^{3}} & \frac{35}{2^{14}} & \frac{315}{2^{7}} \\ & \frac{13}{2^{2}} & \frac{69}{2^{14}} & \frac{165}{2^{5}} & \frac{1505}{2^{8}} \\ & \frac{13}{2^{2}} & \frac{69}{2^{14}} & \frac{165}{2^{7}} & \frac{1505}{2^{8}} \\ & & \frac{13}{2^{2}} & \frac{10005}{2^{10}} \\ & & 2957/2^{8} & \frac{28257}{2^{11}} \\ & & & 288473/2^{14} \end{cases}$$
(78)



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